METHOD OF CONSTRUCTING AND INVESTIGATING STABILITY OF PERIODIC MOTIONS OF AUTONOMOUS HAMILTONIAN SYSTEMS

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An algorithm is proposed for the construction and investigation of orbital stability of Liapunov motions close to the equilibrium position of autonomous Hamiltonian systems with many degrees of freedom.

1. Let us consider the autonomous system of Hamilton's differential equations with N+1 degrees of freedom. We select the generalized coordinates and moments so that the origin of phase space coordinates is the equilibrium position, and represent the Hamiltonian, which is analytic in the neighborhood of the equilibrium position, in the form of series

$$H = H_2 + \ldots + H_m + \ldots$$
 (1.1)

where H_m is a homogeneous polynomial of power m in coordinates and momenta whose coefficients depend on the problem parameters.

First, we consider the system with Hamiltonian $\ensuremath{H_2}$, representing the determining equation in the form

$$D(\sigma) = |a_{ik} - \sigma \delta_{ik}| = 0$$
 (1.2)

where δ_{jk} $(j, k = 0, 1, \ldots N)$ is the Kronecker delta and a_{jk} are coefficients of the linear system of differential equations.

If Eq. (1.2) has a pair of pure imaginary roots of the form $\sigma_0=\pm i\lambda_0$ and among its other roots there are none equal to σ_0 , then the coordinates and momenta q_0,q_j,p_0 , and p_j $(j=1,\ldots,N)$ can be chosen so that function H_2 is of the form

$$H_2 = \frac{1}{2} \lambda_0 \left(q_0^2 + p_0^2 \right) + H_2^{(N)} \tag{1.3}$$

where $H_2^{(N)}$ is a function that depends only on the variables q_j and p_j . It can be checked that the differential equations which correspond to the Hamiltonian (1.3) have the particular solution

$$q_0 = \alpha_0 \sin(\lambda_0 t + \beta_0), \quad p_0 = \alpha_0 \cos(\lambda_0 t + \beta_0)$$

 $q_j = p_j = 0 \quad (j = 1, ..., N)$
(1.4)

where α_0 and β_0 are arbitrary constants that depend on initial conditions. Solution (1.4) is periodic of period $T_0 = 2\pi/|\lambda_0|$.

If Eq. (1.2) has no other roots of the form $\pm in_0\lambda_0$ (n_0 is an integer), then according to Liapunov's theorem on the holomorphic integral [1], a periodic motion of period close to T_0 exists in the nonlinear system of differential equations with the Hamiltonian H. The periodic motions constitute a set whose parameter is called the amplitude (or the orbital parameter) of periodic motion ϵ and depends in initial conditions. Liapunov had also proposed a procedure for constructing periodic motions

in the form of series in powers of the orbital parameter ϵ . Since then this problem was extensively developed, and at present many methods of constructing periodic motions are available. A method of constructing periodic motions based on canonical transformations is also proposed here, although the main aim is to develop an algorithm for solving the problem of periodic motion stability in the strictly nonlinear meaning. The derived method of periodic motion construction is adapted to the solution of this basic problem. This paper is a continuation and natural generalization of the earlier investigation by the authors (*).

2. Periodic motions are Liapunov unstable relative to perturbations of coordinates and momenta q_0 and p_0 , since their period depends on initial conditions. The problem of periodic motion orbital stability is, however, interesting.

In that method the value of constant energy is not fixed, and it can be varied in a certain range. Hence no reduction is made to the number of degrees of freedom, as used in the isoenergetic reduction. This makes it possible to investigate the complete neighborhood of periodic motion using canonical transformations, and in the periodic motion neighborhood it is possible to introduce such local coordinates that the Hamiltonian of perturbed motion is of a form similar to the normal form in the neighborhood of the equilibrium position. In this way the orbital stability problem of periodic motion reduces to that of Liapunov stability relative to local coordinates.

In the considered problem the constructive application of the local method may be schematically outlined as the sequence of the following operations:

- 1) determination of the investigated periodic motion in "action-angle" variables;
- 2) introduction of local coordinates in the periodic motion neighborhood, and the determination of the perturbed motion Hamiltonian;
- 3) passing to the new "angle" variable, linear normalization, and obtaining statements about stability in linear approximation;
- 4) reverting to the old independent variable and effecting the nonlinear normalization of the Hamiltonian;
- 5) using the properties of coefficients of the Hamiltonian normal form to determine the orbital stability of the periodic motion.
- 3. When among the roots of the determining equation (1, 2) there is at least one pair of roots with nonzero real part, the considered periodic motion is unstable. In what follows we assume that all roots of Eq. (1, 2) are pure imaginary and that there are among them no equal or zero roots. It can be then considered that the quadratic part of the Hamiltonian (1, 1) is of the form

$$H_2 = \frac{1}{2} \lambda_0 (q_0^2 + p_0^2) + \frac{1}{2} \sum_{j=1}^{N} \lambda_j (q_j^2 + p_j^2)$$
 (3.1)

where $|\lambda_j|$ are frequencies of the linear system with the Hamiltonian.

^{*)} Markeev, A.P. and Sokol'skii, A.G., Investigation of periodic motions close to Lagrange solutions of the limited three-body problem.

Preprint No. 110, Inst. Prikl. Matem. Akad. Nauk SSSR, 1975.

The form of H_m in (1,1) is

$$H_{m} = \sum_{\mathbf{v}_{0}+\ldots+\mu_{N}=m} h_{\mathbf{v}_{0}\mid \mathbf{v}_{0}\mid \mathbf{v}_{1}\mid \mathbf{v}_{1},\ldots,\mathbf{v}_{N}\mid \mathbf{v}_{N}} q_{0}^{\mathbf{v}_{0}} p_{0}^{\mu_{0}} q_{1}^{\mathbf{v}_{1}} p_{1}^{\mu_{1}} \ldots q_{N}^{\mathbf{v}_{N}} p_{N}^{\mu_{N}}$$
(3.2)

For the construction of periodic motions of the nonlinear problem we use the method of canonical transformations, and represent forms (3.2) as

$$H_m = \sum_{\alpha=0}^m H_{\alpha,\beta} \quad (\beta = m - \alpha)$$
 (3.3)

where $H_{\alpha,\beta}$ denotes the totality of term of power α of variables q_0 and p_0 and of power β of remaining variables. We now carry out the nonlinear canonical transformation

$$q_i, p_i \rightarrow q_i^*, p_i^*$$
 $(i = 0, 1, ..., N), (H \rightarrow H^*, H_2 = H_2^*)$ (3.4)

which in all forms (3.3) of the new Hamiltonian H^* would normalize terms $H_{m,0}^*$ and cancel terms $H_{m-1,1}^*$. Such transformation is convergent [3,4].

Transformation (3.4) can, for instance, be obtained using the classic Birkhoff's method [5]. However that method has a number of disadvantages that are particularly evident in the normalization of the Hamiltonian of systems with many degrees of freedom of considerable order relative to coordinates and momenta. Because of this transformation (3.4) and all subsequent normalizing transformations are better effected by the recently developed method of Hori—Deprit which is based on canonical transformations of Lie [6,7]. It should be noted in this connection that an essential part of the proposed method of periodic motion analysis is the use of the described algorithms on a computer. The basic part of the necessary programs of normalization computation was published by the authors earlier (*).

We represent the generating function of transformation (3.4) which depends only on new (or only on old) variables in the form of series

$$T = T_3 + \ldots + T_m + \ldots$$

The operator equation for the determination of coefficients of the generating function and of the new Hamiltonian is of the form

$$D_0 T_m = H_m' - H_m^* \quad (m = 3, 4, ...)$$
 (3.5)

where operator D_0 is defined as follows:

$$D_0 T_m = -\{H_2; T_m\} = -\sum_{i=0}^{N} \left[\frac{\partial H_2}{\partial q_i} \frac{\partial T_m}{\partial p_i} - \frac{\partial H_2}{\partial p_i} \frac{\partial T_m}{\partial q_i} \right]$$

Here and in what follows braces denote the operation of computation of Poisson's braces. The forms of H_{m} in (3.5) are defined in terms of "lower" functions

$$H_{\alpha}, H_{\beta}^*, T_{\beta} (\alpha = 2, ..., m; \beta = 3, ..., m - 1)$$
 (3.6)

^{*)} Markeev, A. P. and Sokol*skii, A. G., Certain computational algorithms for normalizing Hamiltonian systems. Preprint No. 31, Inst. Prikl. Matem. Akad. Nauk SSSR, 1976.

using Poisson's braces $\{F_{\alpha}; T_{\beta}\} = D_{\beta}F_{\alpha}$, where functions $F_{\dot{\alpha}}$ are expressed in turn in terms of functions (3.6) using similar Poisson's braces. For example,

$$H_{3}' = H_{3}, \quad H_{4}' = H_{4} + \frac{1}{2} D_{3} (H_{3} + H_{3}^{*})$$

$$H_{5}' = H_{5} + \frac{1}{2} D_{3} [H_{4} + H_{4}^{*} + \frac{1}{6} D_{3} (H_{3} - H_{3}^{*})] + \frac{1}{2} D_{4} (H_{3} + H_{3}^{*})$$
(3.7)

If in the course of solving a specific problem it is necessary to consider in the Hamiltonian H terms H_m of order higher than the fifth, then $H_{m'}$ $(m \geqslant 6)$ may be calculated by the formulas proposed by the authors (see footnote on p. 53).

Equation (3.5) decomposes into groups that correspond to terms $H_{\alpha,\beta}$ in formula (3.3). Hence such terms may be normalized independently of each other. Denominators of the form

$$d_{\alpha,\beta} = |\lambda_0| (v_0 - \mu_0) + \sum_{j=1}^N |\lambda_j| (v_j - \mu_j)$$

$$v_0 + \mu_0 = \alpha, \quad \sum_{j=1}^N (v_j + \mu_j) = \beta$$
(3.8)

appear in the normalization of terms $H_{\alpha,\beta}$ in the form (3.2).

Assume that the Liapunov condition of existence of periodic motions of period $2\pi/|\lambda_0|$ is satisfied, i.e. that the problem parameters are such that for all integral n_0 the inequalities

$$\lambda_j \neq n_0 \lambda_0 \quad (j = 1, \dots N)$$
 (3.9)

are satisfied.

Then taking into account (3.9) and that parameters λ_i ($i=0,1,\ldots,N$) are nonzero, we find that in all forms of H_m^* in the new Hamiltonian terms of form $H_{m-1,1}^*$ can be completely cancelled, since denominators $d_{x,\beta}$ which correspond to these terms do not vanish. This also implies that in forms of odd power (i. e. in H_{2m-1}^*) the terms $H_{2m-1,0}^*$ can be cancelled.

In forms of even powers these terms can be normalized and represented in the form

$$H_{2m,0}^* = b_{2m} 2^{-m} (q_0^{*2} + p_0^{*2})^m$$
 (3.10)

where b_{2m} $(2m=4,\,6,\,\ldots)$ depend only on parameters of the problem.

When Eqs. (3.5) have been solved for all m and the related terms of expansion of the generating function determined, the obtained transformation (3.4) is of the form

$$q_{i} = q_{i}^{*} + \sum_{k=1}^{\infty} \frac{1}{k!} D^{k} q_{i}^{*}, \quad |p_{i} = p_{i}^{*} + \sum_{k=1}^{\infty} \frac{1}{k!} D^{k} p_{i}^{*} \quad (i = 0, 1, ..., N)$$

$$D = \sum_{m=3}^{\infty} D_{m}, \quad D^{0} f \equiv f, \quad D^{1} f \equiv D f = \{f; T\}, ...,$$

$$D^{k+1} f = D \left(D^{k} f\right), ...$$

where D is the differential Lie operator and f is an arbitrary function of variables q_i^* and p_i^* (or q_i and p_i).

Let transformation (3,4) have been carried out. Then, retaining previous notation (without asterisks) for the coordinates, momenta and the Hamiltonian, we find that in the new Hamiltonian H the totality of variables q_j and p_j $(j = 1, \ldots, N)$ is of a power not lower than the second. This means that the equations of motion admit particular solutions that correspond to Liapunov's periodic motions for which $p_i = 0$ and the variation of q_0 and p_0 is defined by the equations

$$\frac{dq_0}{dt} = \frac{\partial H}{\partial p_0} = \left[\lambda_0 + \sum_{m=2}^{\infty} m 2^{1-m} b_{2m} (q_0^2 + p_0^2)^{m-1}\right] p_0$$
 (3.12)

$$\frac{dp_0}{dt} = -\frac{\partial H}{\partial q_0} = -\left[\lambda_0 + \sum_{m=2}^{\infty} m2^{1-m} b_{2m} (q_0^2 + p_0^2)^{m-1}\right] q_0$$

In variables action I -angle W, related to q_0 and p_0 by the indicated below formulas Eqs. (3.12), are of the form

$$\frac{dI}{dt} = 0, \quad \frac{dW}{dt} = \lambda_0 + \sum_{m=2}^{\infty} m b_{2m} I^{m-1}$$

$$a = \sqrt{2I} \sin W \qquad n = \sqrt{2I} \cos W$$
(3.13)

$$q_0 = \sqrt{2I} \sin W, \quad p_0 = \sqrt{2I} \cos W$$

The solution of these equations is

$$I = I_0 = \text{const}, \quad W = \Omega_0 (I_0) (t - t_0) + W_0$$

$$\Omega_0 = \lambda_0 + \sum_{m=2}^{\infty} m b_{2m} I_0^{m-1}, \quad T_0 = \frac{2\pi}{|\Omega_0|}$$
(3.14)

where Ω_0 is the frequency and T_0 the period of periodic motion. It will be seen that when $I_0 \to 0$ the motion period tends to $2\pi/|\lambda_0|$.

4. Let us investigate the stability of periodic motion (3.14) relative to perturbations of periodic motion frequency (or what is the same, relative to the perturbation of action variable I_0 of unperturbed periodic motion) and to perturbations of q_j and p_j , $(j=1,\ldots,N)$. Let $\varepsilon=\sqrt{2I_0}$ be a small but finite quantity (only small periodic motions are considered). Let I be the variable of the perturbed motion action related to I_0 by the formula

$$I = \frac{1}{2}\varepsilon^2 + r_0 \tag{4.1}$$

where r_0 is the perturbation of the variable action. The sign of quantity r_0 is bitrary, and q_j and p_j are quantities of the first and r_0 of the second order smallness, and in their meaning all these quantities, unlike e, are infinitely small.

Using (4.1) we define the Cartesian coordinates q_0 and p_0 of perturbed motion in terms of r_0 and ε as follows:

$$q_0 = \sqrt{2I}\sin W, \quad p_0 = \sqrt{2I}\cos W$$
 (4.2)

$$\sqrt{2I} = \varepsilon \left\{ 1 + \frac{r_0}{\varepsilon^2} + \sum_{m=1}^{\infty} \left(-1 \right)^m \frac{(2m-1)!!}{(m+1)!} \left(\frac{r_0}{\varepsilon^2} \right)^{m+1} \right\} =$$
 $\varepsilon + \frac{r_0}{\varepsilon} - \frac{r_0^2}{2\varepsilon^3} + O\left(r_0^3\right)$

Frequency of the considered periodic motion (3.14) in terms of & is of the form

$$\Omega_0 = \sum_{m=0}^{\infty} \Omega_0^{(m)} \varepsilon^m = \lambda_0 + c_{00} \varepsilon^2 + O(\varepsilon^4)
\Omega_0^{(0)} = \lambda_0, \quad \Omega_0^{(2m-1)} = 0, \quad \Omega_0^{(2m)} = (m+1) 2^{-m} b_{2m+2}$$
(4.3)

where c_{00} denote the quantity b_4 from (3.10).

Substituting into the Hamiltonian the quantities (4, 2) and collecting terms of like order relative to q_j , p_j and $\sqrt{|r_0|}$ for the Hamiltonian of perturbed motion we obtain

$$K = K_{2} + K_{3} + K_{4} + \dots$$

$$K_{2} = \Omega_{0}r_{0} + \frac{1}{2} \sum_{j=1}^{N} \lambda_{j} (q_{j}^{2} + p_{j}^{2}) + \sum_{m=1}^{\infty} H_{m,2}^{2}, \quad K_{3} = \sum_{m=0}^{\infty} H_{m,3}^{2}$$

$$K_{4} = B_{00}r_{0}^{2} + \left[\frac{1}{\varepsilon^{2}} \sum_{m=1}^{\infty} \overline{H}_{m,2}^{2}\right] r_{0} + \sum_{m=0}^{\infty} H_{m,4}^{2}$$

$$\overline{H}_{m,2} = \left(q_{0} \frac{\partial}{\partial q_{0}} + p_{0} \frac{\partial}{\partial p_{0}}\right) H_{m,2}$$

$$B_{00} = c_{00} + \sum_{m=1}^{\infty} B_{00}^{(2m)} \varepsilon^{2m}, \quad B_{00}^{(2m)} = (m+1)(m+2) 2^{-(m+1)} b_{2m+4}$$

where the superscript \wedge indicates that the expressions $\varepsilon \sin W$ and $\varepsilon \cos W$ are to be substituted for q_0 and p_0 in the corresponding forms.

Hamiltonian (4.4) is of the period 2π relative to the variable W. Dots in (4.4) denote terms whose order of smallness relative to perturbations is not lower than the sixth. The Hamiltonian of perturbed motion depends first of all on the problem input parameters U and, secondly on parameter ε which defines the periodic motion amplitude (3.14). Note that in the problem of stability the dependence on the initial instant of time t_0 is unimportant.

The investigation of stability in a specific mechanical system shows that in the normalization process the values of parameters $\,U\,$ for which resonance of the first (Liapunov's condition of periodic motion existence), second (generating point of parametric resonance regions), of the third and fourth orders (generating points for the related resonance surfaces) are possible. The general form of such resonance representation is

$$\sum_{j=1}^{N} n_{j} \lambda_{j} = n_{0} \lambda_{0} \qquad \left(\sum_{j=1}^{N} |n_{j}| = n \right)$$
 (4.5)

where n is the order of resonance and n_0 is an arbitrary integer.

When deciding which of resonances are to be taken into account for a complete investigation of stability in multidimensional Hamiltonian systems in the case of a

specific problem the following two remarks must be taken into consideration:

- a) the dependence of the frequency of a system linearized in the equilibrium position neighborhood on parameters U is known, hence in that region of parameter variation the latter are subjected to certain constraints and, consequently, from among all resonances (4.5) only those that are theoretically possible are to be selected;
- b) the structure of the Hamiltonian is often such that some of the theoretically possible resonance do not appear in the course of normalization (i.e. they do not result in the appearance of zero denominators in the generating function), which makes their consideration pointless.

These two remarks considerably lighten the investigation of cases resonance in a specific problem.

5. Let us investigate the stability of a linear system with Hamiltonian K_2 (see (4.4)).

First of all note that the values of parameters U which satisfy relationships (4.5), where $n\!=\!2$ and n_0 is an arbitrary nonzero integer, are generating parameters for the instability region (region of parametric resonance) in the space of parameters U and ε . Without loss of generality such resonances can be represented in the form

$$n_1\lambda_1 + n_2\lambda_2 = n_0\lambda_0 \quad (n_1 \geqslant 0)$$
 (5.1)

where $n_1+\mid n_2\mid=2$. For $n_1=n_2=1$ resonances (5.1) are called parametric resonances of the combinative type, while for $n_1=2$ and $n_2=0$ (or $n_1=0$ and $n_2=2$) they are of the basic type. Resonances (5.1) for which $n_1n_2<0$ do not generate instability regions [8]. We shall, therefore, assume that the numbers n_1 and n_2 are positive.

We shall describe the procedure of normalization of the quadratic part of the perturbed motion Hamiltonian, and represent function K_2 as

$$K_{2} = \Omega_{0} [r_{0} + G_{2} (q_{j}, p_{j}, W)]$$

$$G_{2} = \sum_{m=0}^{\infty} G_{m,2} \varepsilon^{m}, \quad G_{0,2} = \frac{1}{2} \sum_{j=1}^{N} \frac{\lambda_{j}}{\lambda_{0}} (q_{j}^{2} + p_{j}^{2})$$

$$G_{m,2} = \frac{1}{\lambda_{0}} F_{m,2} - \frac{1}{\lambda_{0}} \sum_{k=1}^{m} G_{m-k,2} \Omega_{0}^{(k)}, \quad F_{m,2} = \hat{H_{m,2}} \varepsilon^{-m} \quad (m=1,2,...)$$

Functions $G_{m,2}$ (and also $F_{m,2}$) are of zero order relative to ε , are 2π -periodic functions of W, and are expressed by finite series of sines and cosines of integral multiplicities of W, and their maximum multiplicity does not exceed m.

To reduce function (5, 2) to the normal form it is necessary first to normalize it with respect to variables q_j and p_j $(j=1,\ldots,N)$. For this we pass to the new independent variable W. Then the Hamiltonian that defines the variation of variables q_j and p_j is represented by function G_2 which corresponds to the nonautonomous canonical system with N degrees of freedom.

Normalization of the Hamiltonian G_2 can be carried out by conventional methods, for example, using an algorithm similar to those of Birkhoff or Deprit—Hori. At each step of $G_{m,2}$ normalization it is necessary to solve systems of linear differ-

ential equations with periodic coefficients. However in the considered case these equations have a small parameter $\ \epsilon$ and the "time" W in the right-hand sides of differential equations appears in a special form. Hence the normalization of the nonautonomous canonical system with Hamiltonian G_2 can be reduced to the normalization of an autonomous system (but with N+1 degrees of freedom), i.e. to the solution of a system of algebraic equations.

First of all note that for normalizing function G_2 we use the operator equation (3.5) in the form

$$\Delta_{0}T_{m,2} = G'_{m,2} - G'_{m,2}$$

$$\Delta_{0} = D_{0} - \frac{\partial}{\partial W}, \quad D_{0}T_{m,2} = -\{G_{0,2}; T_{m,2}\}$$
(5.3)

where $T_{m,2}$ and $G_{m,2}^*$ are terms of power m of the expansion of the generating function of the sought transformation T_2 and of the new Hamiltonian G_2^* in series in the small parameter. Functions $G_{m,2}^{'}$ are calculated by formulas similar to (3.7). The process of solving Eqs. (5.3) can be, however, represented in a somewhat different form.

We introduce imaginary variables q_W and p_W defined by formulas

$$q_W = \varepsilon \sin W, \quad p_W = \varepsilon \cos W$$
 (5.4)

After this substitution the time W does not explicitly appear in the Hamiltonian G_2 , since ε and W appear in function K_2 in (4.3) only in the form of combinations (5.4) and ε^2 in frequency (4.3) can be replaced by the expression $q_W^2 + p_W^2$. The obtained Hamiltonian is of the form

$$L = L_2 + \ldots + L_m + \ldots \tag{5.5}$$

$$L_{m} = G_{m-2, 2}^{\vee} \varepsilon^{m-2} = \sum_{\nu_{0} + \dots + \mu_{N} = m} g_{\nu_{0} \mu_{0} \nu_{1} \mu_{1} \dots \nu_{N} \mu_{N}} q_{W}^{\nu_{0}} p_{W}^{\mu_{0}} \dots q_{N}^{\nu_{N}} p_{N}^{\mu_{N}}$$
(5.6)

where the superscript $\check{}$ indicates that in functions marked by it $\mathfrak E$ and W are eliminated by the substitution (5.4); in (5.6) $m \geqslant 3$ and function L_2 is determined below. Note that the effect of operators with superscripts $\check{}$ and $\hat{}$ in Sect. 4 is opposite, hence function $G_{m,2}^{\check{}}$ can be directly obtained from functions $H_{m,2}$ in the second of formulas (4.5). For this it is only necessary to carry out in functions $H_{m,2}$ the formal substitution $q_0 \to q_W$ and $p_0 \to p_W$ and use the last three of formulas (5.2).

Using the rule of composite function differentiation the operator Δ_0 can be presented in the form

$$\Delta_0 = D_0 - \left[p_W \frac{\partial}{\partial q_W} - q_W \frac{\partial}{\partial P_W} \right]$$

and the operation equation (5.3) then becomes

$$\Delta_0 S_{m-2, 2} = L_{m'} - L_{m}^*$$

$$\Delta_0 S_{m-2, 2} = -\{L_2; S_{m-2, 2}\}$$

$$L_2 = \frac{1}{2}(q_W^2 + p_W^2) + \frac{1}{2}\sum_{i=1}^N \frac{\lambda_i}{\lambda_0}(q_i^2 + p_i^2)$$

Functions L_{m} are calculated by the lowest functions using formulas (3.7) where operators D_{m} are to be replaced by operators $D_{m-2,2}$, whose action on an arbitrary function of variables q_{W} , p_{W} , q_{j} , and p_{j} (j=1,...,N) is defined by the relationships

$$D_{m-2,2}F = \{F; S_{m-2,2}\} = \sum_{j=1}^{N} \left[\frac{\partial F}{\partial q_j} \frac{\partial S_{m-2,2}}{\partial p_j} - \frac{\partial F}{\partial p_j} \frac{\partial S_{m-2,2}}{\partial q_j} \right]$$
(5.7)

It is thus possible to normalize instead of the nonautonomous—system with Hamiltonian G_2 —the autonomous system (but with the number of degrees of freedom increased by one) with Hamiltonian (5.5). It is assumed that in this case the normalization procedure differs from that of normalization—of an autonomous system with N+1 degrees of freedom only by that in calculating Poissons braces in the quantities of (3.7) and in deriving the explicit form of (3.11) differentiation is carried—not with respect to all variables q_w , p_w , q_j , and p_j ($j=1,\ldots,N$) but, in accordance with (5.7), only with respect to variables q_j and p_j .

Since variables q_j and p_j appear in Hamiltonian (5.5) in quadratic form (in (5.6) ($v_1 + \mu_1 + \ldots + v_N + \mu_N = 2$)), only resonances of the form (4.5) with n=2 can impede normalization.

In the absence of any such resonances Hamiltonian (5.5) can be reduced to the following normal form:

$$L^* = r_W + \sum_{j=1}^{N} \left\{ \frac{\lambda_j}{\lambda_0} + \sum_{m=1}^{\infty} a_{2m,j} r_W^m \right\} r_j$$

$$q^* = \sqrt{2r} \sin \varphi, \quad p^* = \sqrt{2r} \cos \varphi$$
(5.8)

Function (5.8) is presented in "polar" coordinates r_W , ϕ_W , r_j , and ϕ_j related to the new variables q_W^* , p_W^* , q_j^* , and p_j^* by formulas of the indicated form. Parameter $a_{2m,j}$ depends on parameters U of the problem.

If the resonance relationship (5.1) is satisfied, the normal form of Hamiltonian (5.5) is

$$\begin{split} L^{*} &= r_{W} + \sum_{j=1}^{N} \left\{ \frac{\lambda_{i}}{\lambda_{0}} + \sum_{m=1}^{n_{0}'} a_{2m, j} r_{W}^{m} \right\} r_{j} + \\ & a \sqrt{r_{W}^{\lfloor n_{0} \rfloor} r_{1}^{n_{1}} r_{2}^{n_{2}}} \sin \left(n_{1} \varphi_{1} + n_{2} \varphi_{2} - n_{0} \varphi_{W} \right) + O\left(r_{W}^{\gamma} \right) \\ \gamma &= \frac{\mid n_{0} \mid + 1}{2}, \quad n_{0}' = \left[\frac{1}{2} \mid n_{0} \mid \right] \end{split}$$

where parameter a also depends on parameters U and brackets indicate the taking of the integral part of a number.

When ε (i.e. q_W and p_W) are fairly small the transformation

$$q_W, p_W, q_i, p_i \rightarrow q_W^*, p_W^*, q_i^*, p_i^* \quad (i = 1, ..., N)$$
 (5.9)

is convergent for q_W and p_W . It is carried out using formulas similar to (3.11) in which again operators $D_{m-2,2}$ from (5.7) are to be substituted for operators D_m . The form of these operators clearly shows, as expected, that the imaginary variables q_W and p_W are not affected by such normalization.

Now let us carry out the transformation inverse of (5.4) and revert to the old independent variable t. If, in addition, we specify the transformation r_0 , $W \to r_0^*$, W^* by formulas

$$r_0 = r_0^* + \partial S_2 / \partial W, \quad W = W^*$$
 (5.10)

then together with (5.9) we obtain the canonical transformation which normalizes function (5.2) with respect to all variables. In the nonresonance case the normal form of that function is (previous notation is retained for r_0)

$$K_2^* = \Omega_0 r_0 + \sum_{j=1}^N \Omega_j r_j$$
 (5.11)

and in the resonance case it is

$$K_2^* = \Omega_0 r_0 + \sum_{j=1}^N \Omega_j' r_j + A \epsilon^{|n_0|} \sqrt{r_1^{n_1} r_2^{n_2}} \sin(n_1 \varphi_1 + n_2 \varphi_2 - n_0 W) + (5.12)$$

The following notation is used in these formulas:

$$\Omega_{j} = \sum_{m=0}^{\infty} \Omega_{j}^{(m)} \varepsilon^{m}, \quad \Omega_{j}' = \sum_{m=0}^{2n_{0}'} \Omega_{j}^{(m)} \varepsilon^{m}, \quad A = a\lambda_{0} 2^{-|n_{0}|/2}
\Omega_{j}^{(0)} = \lambda_{j}, \quad \Omega_{j}^{(2m-1)} = 0, \quad \Omega_{j}^{(2m)} = \frac{\lambda_{j}}{\lambda_{0}} \Omega_{0}^{(2m)} + \sum_{k=1}^{m} 2^{-k} a_{2k, j} \Omega_{0}^{(2m-2k)}$$

Let us consider the case of parametric resonance. Regions of parametric resonance (instability regions) issue for small ϵ from surfaces for which in the region of variation of the problem parameters U relationship (5.1) is satisfied. According to [8] on the surfaces which bound these regions in space of parameters U and ϵ the following relationships are valid:

$$|n_{1}(\lambda_{0}\Omega_{1} - \lambda_{1}\Omega_{0}) + n_{2}(\lambda_{0}\Omega_{2} - \lambda_{2}\Omega_{0})| = |\lambda_{0}A| \sqrt{n_{1}^{n_{1}}n_{2}^{n_{2}}} \varepsilon^{|n_{0}|}$$
 (5.13)

When the right-hand side of the last relationship is greater than the left-hand, the periodic motion is unstable, and when it is smaller, we have stability in the first approximation.

Equations (5.13) of two surfaces in the space of parameters U and ε may be sought in the form of series in ε , using series expansions of Ω_0 , Ω_1 , Ω_2 .

6. If parameters U and & of the problem are such that the considered periodic motion is stable in linear approximation, then by normalizing the linear system using the method described in Section 5, Hamiltonian (4.4) can be reduced to the form

$$K^* = K_2^* + K_3^* + K_4^* + \dots$$

$$K_3^* = \sum_{m=0}^{\infty} \hat{K_{m,3}}$$
(6.1)

$$K_4^* = B_{00}r_0^2 + \left[\frac{1}{\epsilon^2}\sum_{m=1}^{\infty}\hat{K_{m,2}}\right]r_0 + \sum_{m=0}^{\infty}\hat{K_{m,4}}$$

where K_2^* is of the form (5.11) and functions $K_{m,i}^{\hat{}}$ ($m=0,1,\ldots;i=2,3,4$) are of order m relative to ϵ (i.e. relative to the imaginary variables q_W and p_W in Sect. 5) and of order i relative to q_j and p_j ($j=1,\ldots,N$). These functions are readily calculated by formulas similar to (3.7). For instance,

$$K_{0,3} = H_{0,3}, K_{1,3} = H_{1,3} + D_{1,2}H_{0,3}$$

$$K_{2,3} = H_{2,3} + D_{1,2}H_{1,3} + D_{2,2}H_{0,3}, K_{1,2} = \overline{H}_{1,2}$$

$$K_{2,2} = \overline{H}_{2,2} + D_{1,2}\overline{H}_{1,2}, K_{0,4} = H_{0,4} + \frac{1}{\varepsilon^2}H_{1,2} \left[q_W \frac{\partial}{\partial p_W} - p_W \frac{\partial}{\partial q_W} \right] S_{1,2}$$

$$(6.2)$$

Elucidation of the question of stability in the strict (nonlinear) meaning, requires extension of the normalization process of the Hamiltonian of perturbed motion.

Normalization of Hamiltonian (6.1) can be impeded by resonances

$$\sum_{j=1}^{N} n_{j} \Omega_{j} = n_{0} \Omega_{0} \quad \left(\sum_{j=1}^{N} |n_{j}| = n = 3, 4 \right)$$
 (6.3)

In the region of variation of parameters U and ϵ formulas (6.3) are equations of resonance surfaces of the third and fourth order and are derived similarly to the boundaries of parametric resonance regions in Sect. 5. Parameters U that satisfy relationships (4.5) are generating parameters for such surfaces.

Let us, first, consider the values of parameters U and ε which are not associated with third and fourth order resonance surfaces. In that case form K_3^* in Hamiltonian (6.1) can be completely eliminated by using the method described in Sect. 5. Normalization of fourth order terms consists of the following three independent stages.

- a) Normalization of terms proportional to r_0^2 . These terms are already normalized.
- b) Normalization of terms proportional to r_0 . It can be shown that the normalization of these terms reduces to the averaging of function $K_{1,2}+K_{2,2}+\dots$ with respect to rapid phases of motion determined by Hamiltonian (5.11). Note that for n=2 resonances (6.2) do not impede the normalization of these terms, since they appear only at the boundaries of parametric resonance regions and, consequently have been already taken into account in linear normalization.
- c) Normalization of terms independent of r_0 . This normalization stage is similar to linear normalization procedure.

As the result, Hamiltonian (6.1) of perturbed motion can in the nonresonance case be reduced to the following normal form (previous notation is used for variables):

$$K = K_2 + K_4 + K^*$$

$$K_2(r_0, r_1, \dots, r_N) = \sum_{i=0}^{N} \Omega_i r_i, \quad K_4(r_0, r_1, \dots, r_N) = \sum_{0 \le i \le j \le N} B_{ij} r_j r_j$$
(6.4)

where the expansion of coefficients of form K_4 in series in ε is similar to that of B_{00} in (4.4), and K^* is a 2π -periodic function of angle variables $W, \varphi_1, \ldots, \varphi_N$; its order relative to r_i is not less than $^{5}/_{2}$.

In the case of the third order resonance the normal form is

$$K = K_2 + A \varepsilon^{|n_0|} \sqrt{r_1^{|n_1|} \dots r_N^{|n_N|}} \sin(n_1 \varphi_1 + \dots + n_N \varphi_N - n_0 W) + K^*$$
 (6.5)

and in that of the fourth order resonance it is

$$K = K_2 + K_4 + A \varepsilon^{|n_0|} \sqrt{r_1^{|n_1|} \dots r_N^{|n_N|}} \sin(n_1 \varphi_1 + \dots + n_N \varphi_N - n_0 W) + K^*$$
 (6.6)

The order of function K^* relative to ε in (6.5) and (6.6) is not lower than $|n_0|+1$, and the quantities B_{ij} in (6.6) are determined with that accuracy.

7. Thus for determining the stability of a periodic motion it is only necessary to calculate the coefficients of one of the normal forms (6,4)- (6,6) and apply the stability criteria from [4,9-14]. Comprehensive results may be obtained in this manner for systems with two degrees of freedom (N=1).

If third or fourth order resonances (6.3) appear in a multidimensional Hamiltonian system and in the array of numbers n_1, \ldots, n_N at least one change of sign takes place, the periodic motion is formally stable [9], i.e. it is stable in any approximation.

When third order resonance (6.3) is present and in (6.5) $A \neq 0$, the periodic motion is unstable [10,11]. If A = 0, the question of stability is not resolved by terms of that order.

If fourth order resonance (6, 3) is present, and in the normal form (6, 6)

$$|K_4(-n_0, n_1, ..., n_N)| < |A| \sqrt{n_1^{n_1} ... n_N^{n_N}} \varepsilon^{|n_0|}$$
 (7.1)

the periodic motion is unstable [10, 11]. With the opposite sign in this inequality in the case of two-frequency Hamiltonian systems we have stability [12], and in the multi-dimensional case stability is shown in the last (fourth) approximation [11].

In the nonresonance case of systems with two degress of freedom the question of stability is resolved by the Amold—Moser theorem, viz. if (in notation of (6.4) and (4.4))

$$D \neq 0 \tag{7.2}$$

$$D = K_4 (\Omega_1, -\Omega_0, 0, \ldots, 0) = c_{00}\lambda_1^2 - c_{01}\lambda_0\lambda_1 + c_{11}\lambda_0^2 + O(\epsilon^2)$$
 (7.3)

the periodic motion is stable [4, 13].

The state of development of the theory of Hamiltonian systems does not provide means for obtaining a similarly complete result in the multidimensional case. It is only possible to make the following statement.

If for
$$r_0 = r_1 = ... = r_N = 0$$
 the determinants

$$D_{3} = \det \left\| \frac{\partial^{2} K_{4}}{\partial r_{i} \partial r_{j}} \right\|, \quad D_{4} = \det \left\| \begin{array}{cc} \frac{\partial^{2} K_{4}}{\partial r_{i} \partial r_{j}} & \frac{\partial K_{2}}{\partial r_{i}} \\ \frac{\partial K_{2}}{\partial r_{j}} & 0 \end{array} \right\|$$
(7.4)

do not simultaneously vanish, stability is present for the majority (in the meaning of the Lebesgue measure) of initial conditions [13].

It is also possible to consider the problem of formal stability of periodic motions. In the considered case the sufficient condition of formal stability reduces (see [14] and the footnote on p. 52) to the check of incompatibility of the system of equations (relative to r_0, r_1, \ldots, r_N)

$$K_2 = 0, K_4 = 0$$
 (7.5)

in the region $r_1 \ge 0, \ldots, r_N \ge 0$ (note that by definition (4.1) the sign of parameter r_0 is arbitrary).

In determinants (7.4) and Eqs. (7.5) it is evidently, reasonable to take into account only the principal terms of expansions of Ω_i , and B_{ij} in (6.4) (see also (7.3)). This means that for solving the question of stability in the nonresonance case, it is possible as a rule, to restrict the analysis to and including terms H_4 of the input Hamiltonian (1.1).

REFERENCES

- 1. Liapunov, A.M., The general problem of motion stability. Collected Works.

 Moscow Leningrad, Izd. Akad. Nauk SSSR, 1956.
- 2. Briuno, A.D., Instability in Hamiltonian system and asteroid distribution. Matem. Sb., Vol. 83, No. 2, 1970.
- 3. Siegel, K.L., Lectures on Celestial Mechanics. Moscow, Izd. Inostr. Lit., 1959.
- 4. Moser, J., Lectures on Hamiltonian Systems. Providence, New Jersey. Mem. Amer. Math. Soc. No. 81, 1968.
- 5. Birkhoff, G.D., Dynamical Systems. New York, Amer. Math. Soc. No. 9, Providence, 1927.
- 6. Hori, G.I., Theory of general perturbations with unspecified canonical variables. J. Japan Astron. Soc. Vol. 18, No. 4, 1966.
- 7. Deprit, A., Canonical transformations depending on a small parameter. Celest Mech., Vol. 1, No. 1, 1969.
- 8. Iakubovich, V.A. and Starzhinskii, V.M., Linear Differential Equations with Periodic Coefficients and their Applications. Moscow, "Nauka", 1972.
- Moser, J., New aspects in the theory of stability of Hamiltonian systems.
 Communs. Pure and Appl. Math., Vol. 11, No. 1, 1958.
- 10. Markeev, A.P., Stability of a canonical system with two degrees of freedom

- in the presence of resonance. PMM, Vol. 32, No. 4, 1968.
- 11. Khazin, L.G., On the stability of Hamiltonian systems in the presence of resonances. PMM, Vol. 35, No. 3, 1971.
- 12. Markeev, A.P., On the problem of stability of equilibrium positions of Hamiltonian systems. PMM, Vol. 34, No. 6, 1970.
- 13. Arnol^od, V.I., Small denominators and the problem of stability of motion in classical and celestial mechanics. Uspekhi Matem. Nauk, Vol. 18, No. 6, 1963.
- 14. Markeev, A.P., On the stability problem for the Lagrange solutions of the restricted three-body problem. PMM, Vol. 37, No. 4, 1973.

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